CAMBRIDGE CS

Compressed sensing reconstruction for multidimensional NMR spectroscopy

User manual

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Chapter 1: Overview

Methods for improving the time-efficiency of NMR data acquisition have been popular since the early days of multidimensional NMR [1]. Such methods have focussed around recording a subset of the full data matrix (non-uniform sampling, NUS) and coupled with a variety of processing methods due to the restriction of the Fourier transform to using regularly sampled data [1–7]. NUS approaches are typically used to improve resolution, sensitivity or reduce the recording time for spectra, or a combination of all three [4,8]. Compressed sensing (CS) was pioneered in the field of information theory [9–11] and quickly found applications in a number of fields including MRI [12] and more recently NMR as a new method for processing NUS data [13–16]. CS has been demonstrated to be a powerful method to reconstruct undersampled spectra with high fidelity, even for weak peaks in spectra with a high dynamic range [16]. Consequently CS has rapidly gained popularity in NMR spectroscopy. The aim of this software is to provide a simple means of processing NUS data with a variety of available CS algorithms.

Chapter 2: Installation

The software comes as a single .tar.gz package. In order to simplify installation the package includes miniconda3 containing python3 and relevant libraries. This will hopefully avoid problems with different python/python package versions. Miniconda is subject to the Anaconda End User Licence Agreement see (https://docs.continuum.io/anaconda/eula). The package also contains Azara (see http://www2.ccpn.ac.uk/azara/) which is used to provide postprocessing commands.

The current version has been tested on 64 bit Linux distributions. A separate download is available for 64 bit Mac OS X 10.11 and later. To install:

1. Uncompress and extract the download.

tar zxvf cambridgecs.tar.gz

2. The uncompressed folder *CambridgeCS* contains the following folders: *azara-2.8*, *bin*, *binBuild*, *miniconda3*, *testData*. Change into the *binBuild* directory and run the command:

bash buildCython

This will produce .so files for all the .c files in binBuild.

3. The code is called using the scripts in the *bin* directory. These should be added to the system path e.g. to your *.bashrc* file. cambridgecs calls the reconstruction code, whilst graphPlot calls the simple spectrum viewer independently.

Chapter 3: Quick Start

This section is designed to give a very brief overview of running the code in simple cases, with a particular view to running the example scripts contained in the folder *testData* and to adapting these to other similar situations. Detailed information on running 'Cambridge CS' and the various commands therein is given in chapter ??.

CambridgeCS can be called using the following command:

cambridgecs [OPTION] [Script file]

```
-a --aut
Run in automatic mode. Does not launch GUI and requirements for user input are
suppressed. Requires a script file.
-h --help
Show help text
-g --gui
Launch the GUI. The GUI maybe launched without a script file in which case a script
can be built within the GUI. Alternatively an existing script file maybe loaded
into the GUI and amended or run.
```

To load one of the test scripts (e.g. $550_cs.scr$), change to the relevant directory and run the command:

cambridgecs -g 550_cs.scr

The example chosen is for a 3D H-detected HNCO experiment. The GUI will load the commands contained in the script file. Basic information about the dataset is shown in the 'Data parameters' section. Below this is the 'Pre-processing script' which carries out processing of the direct dimension. Additional commands can be added using the green crosses. Red crosses remove commands. The data reconstruction is carried out using the 'Reconstruction script' section. Here the choice of reconstruction algorithm can be made along with phasing and weighting functions for the indirect dimensions. Finally post processing is specified in the 'Post-processing script'. Typically this involves additional commands e.g. reversing dimensions or adding additional baseline corrections. As with the direct dimension, processing uses Azara (packaged with the code), the dimensions are specified using script_com_n for the nth dimension.

In the 'Data Parameters' and 'Reconstruction Script' sections of the code the dropdown 'Extra Params' menus contain additional, but non-essential commands which may be used or needed in certain situations.

The script can be viewed with the button 'Open script editor'. This loads the 'Script editor' window where the script can also be edited and synced with the GUI. It is hoped that this will help users to become familiar with the scripting style. The script can be saved with the button 'Save file' although the script will also be automatically saved once the code is run.

To run the code, press the button 'Process current script'. A terminal window will appear to show the progress of the reconstruction. At the end of the reconstruction the simple viewer will open automatically (assuming post processing commands are set). This allows quick visual inspection of the reconstruction. Changes can be made in the GUI or the 'Script editor' and the code re-run as necessary. Each run opens a new simple viewer window. GUI windows can be closed with the keys 'q' or 'Esc'.

The *testData* folder contains other examples of 2, 3 and 4D data sets all of which can be reconstructed using CambridgeCS (see section B for details).

Whilst the main use of Cambridge CS is for CS processing of NUS data, it is also possible to process fully-sampled data. This can be done in two ways - either by treating the data as if it is an NUS experiment by providing a suitable points list (nuslist). This is essential if the data is recorded in NUS order i.e. quadrature detection is performed before dimension incrementation. If the experiment is recorded as a standard fully-sampled experiment the 'Uniform sampling flag' can be set in 'Data Parameters'. It is then necessary to specify 'max_points' for each indirect dimension under the 'Reconstruction script' so that the code knows how to correctly reshape the data. An example of this is given in the script 102_ft_full.scr.

Chapter 4: Using Cambridge CS

4.1 Concept

Data processing may be carried out using either a script-based approach, or using the GUI. We recommend that initially the GUI is used in order to become familiar with the scripting approach. The button 'Open Raw Script' opens a text window which will be populated with commands as these are selected in the GUI. Commands may also be typed in the scripting window and synced with the GUI via the 'Sync text' button. The scripts and GUI have four sections:

Data parameters Pre-processing script Reconstruction script Post-processing script

'Data parameters' contains information about the data i.e. file size, location, data type etc. Data parameters must be set for all processing scripts. More details are given in the section 4.2.

'Pre-processing' parameters contains the commands needed for processing the direct (acquisition) dimension of multidimensional experiments. It is assumed that this dimension is always fully-sampled. More details on the pre-processing functions are given in section 4.3. This section is optional if pre-processing is carried out in another software package. Explicit support is available for data pre-processed via the NMRPipe suite (section 4.2.2).

Reconstruction parameters contains the options for NUS data reconstruction of the indirect dimensions either via the FFT, or using a variety of compressed sensing algorithms. The output from this may either be reconstructed time-domain data, or reconstructed frequency-domain data.

'Post-processing'; this section is optional. If post-processing commands are set in the script, processing will be carried out using the Azara software package (W. Boucher, unpublished). Both time-domain or frequency-domain output from the reconstruction may be processed. Alternatively other processing packages may be used. Further details on Azara post-processing are given in section 4.5. Note that the post-processing section must be set for the simple viewer to open automatically.

4.2 Loading data

4.2.1 Data Parameters

Various options and flags are available to specify data import parameters. These are detailed below. Essential options are visible in the GUI. Extra options may be selected from the dropdown menus.

Directory

directory </Path/To/Directory>

Specify the file path to the directory containing the input data. Output data will be written to this directory. See also **Raw data directory** below

Data Input

dataInput </Path/To/Data>

Specify the file path to the input data if this is not the raw ser or fid file (Bruker and Varian respectively).

Raw data directory

rawDirectory <Directory>

Raw data may be stored in a subdirectory within Directory. If Raw data directory is specified, the software will look for raw data in </Path/To/Directory/rawDirectory>. Output data will still be saved in </Path/To/Directory>.

NUS list

nus List <
name>

Specify the name of the list file containing the acquired points in the indirect dimension(s). This may be specified in terms of increments (i.e. starting at 0) or in terms of points (i.e. starting at 1). The file format may be a column vector, where a single point/increment is specified by ndim lines where ndim is the number of dimensions, or as a matrix of the size [npts_indirect, ndim] where npts_indirect is the total number of points in the indirect dimensions. The software detects the format automatically. The NUS list is assumed to be in the experiment directory (either **Directory**, or **Raw data directory**). The default is *vclist* (Bruker convention)

Original NUS list

nusListOriginal <name>

Used in combination with <nusList> to artificially further undersample an experiment. <nusListOriginal> specifies the nusList which describes the recorded data. <nusList> is then used to specify a second list which which specifies a new undersampled schedule. For example, if a fully-sampled experiment is recorded and the user wishes to test the effect of 10% sampling, <nusList> would be set to a 10% sampling list while <nusListOriginal> would be a fully-sampled list. The data is then read-in using <nusListOriginal> and masked according to <nusList>. The output of the reconstruction will be equivalent to the 10% sampled spectrum. If the original experiment is already undersampled, care must be taken to avoid specifying points that are not present in <nusListOriginal>.

Data type

int, float

Used to specify the data type of raw data: either 32-bit integer or 32-bit float. Note that this option is only available for raw (unprocessed data) which is being processed using the pre-processing commands. If data is pre-processed in other software, it is assumed the data-type is 32-bit float.

Endian

big_endian, little_endian

Used to specify the endianess of raw data. Note this option is only available for raw (unprocessed data) which is being processed using the pre-processing commands. If data is pre-processed in other software, it is assumed that the data-type is 32-bit float.

Data input flag

nmrPipe, varian

Used to specify alternative data input formats (NMRPipe or Varian).

Shuffle

shuffle21, shuffle32, shuffle5724

Shuffles the input data ordering by switching the locations of real (r) and imaginary (i) points in the indirect dimensions as shown below:

Dimension	Default order		Shuffled order	
2D	r	1	i	2
	i	2	r	1
3D	rr	1	rr	1
	ir	2	ri	3
	ri	3	ir	2
	ii	4	ii	4
4D	rrr	1	rrr	1
	irr	2	\mathbf{rri}	5
	rir	3	rir	3
	iir	4	rii	7
	rri	5	irr	2
	iri	6	iri	6
	rii	7	iir	4
	iii	8	iii	8

Threads

threads <number>

Specify the number of cpu threads used to reconstruct the data. The default is 1. If set to 0, the maximum number of available cores will be used.

4.2.1.1 Data size

All data is imported as a pseudo-2D.

Direct-dim

```
npts npts <points>
```

The total number of points in the direct dimension (real and imaginary).

block block <block size>

The block size for the direct dimension. This option can be found in the drop-down menu 'Extra Params'.

Indirect-dims

npts npts <points>

The total number of points in all the indirect dimension (real and imaginary) i.e. product of the indirect dimension sizes. ndim ndim <number>

The total number of indirect dimensions, i.e. n-1 for an *n*-dimensional experiment.

block block <block size>

The block size for the indirect dimensions. This option can be found in the drop-down menu 'Extra Params'.

interlace interlace <dimension>

Used to convert P and N type data e.g. from Echo-Antiecho/Rance-Kay data acquisition to cosine/sine type data needed for processing to generate an absoprtion mode lineshape. Note interlacing is only used if a pre-processing script is included. Dimension assumes that n = 1 is the direct (acquisition dimension).

acquired_points acquired_points <number>

Used in cases where the acquisition was not completed to specify the number of data points (real and imaginary) which were recorded in the indirect dimensions as opposed to the total size of the indirect dimensions which should still be specified.

4.2.1.2 RQD-params

More information forthcoming

4.2.2 NMRPipe

Cambridge CS may be be used in conjunction with NMRPipe processing. A typical scenario would be to use NMRPipe for pre- and post-processing with CS reconstruction carried out in the indirect dimensions using the algorithms available in Cambridge CS. In this case the CS processing script should not contain the sections

- # Pre-processing parameters:
- # Azara post-processing parameters

Cambridge CS is able to extract parameter information from NMRPipe binary files, and write out to NMRPipe format for further processing/viewing in the NMRPipe software suite. One significant difference between NMRPipe and Cambridge CS is that Cambridge CS uses the Azara convention of specifying dimensions in numerical order and processing in situ, rather than the NMRPipe convention of transposing the required dimension to the direct dimension for processing. Thus a typical NMRPipe pre-processing script should not include the final transpose statement.

|nmrPipe -fn TP ∖

A sample NMRPipe pre-processing script for a 2D TROSY is shown below:

Conversion from Bruker format to pipe format:

```
<fid.com>
#!/bin/csh -f
#
bruk2pipe -in ./ser -bad 0.0 -DMX -noswap \
         -decim 2000 -dspfvs 20 -grpdly 67.9862518310547 \
-xN
         2048
                    -vN
                             300
                                            \
                             250
-xT
         1024
                    -yT
                                            \
                    -yMODE Echo-AntiEcho \
-xMODE
         Complex
-xSW
         10000
                    -ySW
                             1700.680
                                            /
-xOBS
         800.130
                    -yOBS
                             81.076
                                            ١
-xCAR
         4.708
                    -yCAR
                             115.100
                                            ١
-xLAB
         1H
                     -yLAB
                             15N
                                            ١
-ndim
         2
                    -aq2D
                             States
-out data.fid -ov -verb
```

Processing of the direction dimension, keeping the data in the form xy:

```
nmrPipe -in data.fid
| nmrPipe -fn POLY -time
| nmrPipe -fn SP -off 0.5 -end 0.98 -pow 2 -c 1.0
                                                  \
| nmrPipe -fn ZF -size 4096
| nmrPipe -fn FT
| nmrPipe -fn PS -p0 -173 -p1 180 -di
| nmrPipe -fn POLY -auto -xn 5.0ppm -ord 1
| nmrPipe -fn EXT -xn 5.7ppm -x1 10.95 -sw -verb
-verb -ov -out test.nus
```

Subsequent CS processing requires specification of a header flag <input_flag> with the value 'nmrPipe', as well as the size of the NMRPipe header <head> (the default value is 512 which should work). Other parameters are specified as normal. An example script is shown below:

\

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```
# General properties:
```

<ft1.com>

```
directory
                      /home/user/path/to/pipe/input/file
                      /home/user/path/to/pipe/input/file/test.nus
    dataInput
rawDirectory Bruker_directory_number
    outputFile
                      pipe_output
    experimentNumber
                      1
    nmrPipe
    head
                      512
    direct-dim:
                             1714
        npts
end_script
    indirect-dims:
                             300
        npts
```

```
ndim
                             1
    end_script
# Reconstruction parameters:
    reconstruction_script:
        reconstruction_flag IHT
        stop_flag
                             12
        dim 2:
            zerofill
                             2
            phase
                             -90.0
                                       -360.0
                             60.0
            sinebell
        end_script
    end_script
```

After processing in the indirect dimensions, a small NMRPipe script <\$outputFile.com> is written automatically to convert the data back to a format which can be read by NMRPipe, using the NMRPipe programme bin2pipe as shown below.

```
#!/bin/csh
#
bin2pipe \
   -in /home/user/path/to/pipe/input/file/pipe_output_recon.spc \
   -bad 0.0 -noswap \
-xN 1714 \
-xT 1714 ∖
-xFT Freq \
-xMODE Real \
-xSW 4184.570312 \
-xORIG 4560.469238 \
-xOBS 800.130005 \
-xLAB 1H \
-yN 2048 \
-yT 2048 ∖
-yFT Freq ∖
-yMODE Real \
-ySW 1700.680054 \
-yORIG 8492.844727 \
-yOBS 81.075996 \
-yLAB 1H ∖
-ndim 2 \
InmrPipe -fn TP -auto \
|nmrPipe −fn REV \
|nmrPipe -fn TP -auto \
-out /home/user/path/to/pipe/input/file/pipe_output.pipe.ft2
     -ov -verb
```

The various parameters shown in this script are extracted from the header file of the input NMRPipe file automatically. Further examples of using CambridgeCS in combination with NMRPipe are given in Appendix B

4.3 Pre-processing (direct dimension processing)

Pre-processing of NUS data involves processing the directly acquired dimension which is fully-sampled. This may be done using the software of choice of the user e.g. Bruker Topspin or NMRPipe. Loading such data is described in section 4.2. Within the CS-NMR software, we provide direct-dimension processing modelled on the Azara software but coded in Python (W. Boucher, unpublished). Common processing functions, described below, are available and operate in the same way as described in the Azara help pages (<http://www2.ccpn.ac.uk/azara/azara_docs/process.html>), which may be useful for further information. The direct dimension processing commands should be contained between the flag-lines pre_processing_script: and end_script. An example pre-processing script is provided below:

```
pre_processing_script:
    complex
    brukerGroupPhase
    sinebell
                      60
                      16.4
                                    10000
    gaussian_sw
                              1.1
    zerofill
                      2
    fft
                      7
                             -72
    phase
    real
                      1400
                              2048
    range
end_script
```

The output from the pre-processing script is used directly for further processing, but the data processed in dimension 1 is also written out directly as $*_prepro_f1.spc$ along with $*_prepro_f1.spc.par$, a parameter file containing information on the dimension and block sizes in the Azara format, which may be of use for further processing in other software packages or to analyse the direct dimension further e.g. using the Azara plot2 software. Only information on the number of points and block sizes are correct in the parameter file. Values for spectral width (sw), spectral frequency (sf), reference point (refpt), reference ppm (refppm) and nucleus (nuc) are arbitrary and would need to be edited by the user if correct referencing is needed. Most likely these files will not be used by the average user.

Processing commands available are listed below with a brief describtion of their usage and syntax, taken from the Azara manuel. For more information, see the Azara help section <http://www2.ccpn.ac.uk/azara/azara_docs/process.html>. In the GUI, commands may be selected by clicking on the green cross and selecting the relevant command. Any necessary parameters should be typed in the text boxes. Commands may be removed by clicking the red cross

complex

complex

Specify data is complex.

real

```
real
Specify data is real.
```

conjugate

conjugate

Take the complex conjugate of data. Data must be complex, i.e. the complex flag (above) must previously have been set.

brukerGroupPhase

brukerGroupPhase <GRPDLY>

This does the first order phase correction required for Bruker AV data where $\langle \text{GRPDLY} \rangle$ is given in the Bruker *acqus* file. The phase correction is $360 \times \langle \text{GRPDLY} \rangle$. For Bruker Avance data the parameters DSPFVS and DECIM, given in the Bruker *acqus* file can be converted to a lookup value using the lookup table (see ?? and converted to $\langle \text{GRPDLY} \rangle$) using the formula groupDelay= $0.5 \times \text{lookup/DECIM}$).

If no argument is specified GRPDLY will be read directly from the *acqus* file assuming this is in the experiment directory.

The brukerGroupPhase command applies the following functions:

fftn phase 0 <360.0*GRPDLY> ifftn

where fftn and ifftn are the normalised Fourier and inverse Fourier transforms respectively, phase is as specified below, and upper cuts the upper point of the data to the initial number of points (npts1) minus the group delay value, with int(\$GRPDLY) retaining the integer part of group delay.

$conv_sine$

 $conv_sine < half width >$

This convolves the data with a sine function with the given <half width>, and subtracts the result from the data. Typically this is done on the acquisition fid before anything else is done, in order to remove a water signal that occurs at zero frequency. A typical value with 1024 data points would be conv_sine 8.

decay_sw

decay_sw <line broadening> <spectral width>

This multiplies the data by a decaying exponential $\exp(-\pi LBt)$ where LB is the linebroadening (Hz).

sinebell

sinebell <angle>

This multiplies the data by a sine function with the given $\langle angle \rangle$ (specified in degrees) at point 1, and, if there are n (complex/real) points, with angle 180 degrees at point n+1 (*not* point n).

sinebell 2

sinebell2 < angle >

This multiplies the data by a sine function squared with the given $\langle angle \rangle$ (specified in degrees) at point 1, and, if there are n (complex/real) points, with angle 180 degrees at point n+1 (*not* point n).

gaussian_sw

gaussian_sw line broadening (Hz)> <sharpening factor> <spectral width (Hz)>

This multiplies the data by the gaussian $\exp(a + bt + ct^2)$. If LB is the entered line broadening (in Hz) and s is the sharpening factor, then $a = -\ln 2/s^2$, $b = \pi \times LB$ and $c = -(\pi \times LB \times s)^2/4 \ln 2$. Note that LB must be **positive** for normal use, which is the **opposite** of the Bruker convention. The function converts a Lorentzian line (or multiplet) of width FWWH equal to LB to a Gaussian line of width FWWH equal to $s \times LB$. The maximum value of the multiplying function is 1, which is obtained at the fraction of the acquired time $t/T = 2\ln 2/(LB \times \pi \times T \times s^2)$. The spectral width (SW) is needed by the program in order to translate between point number, x, and the time value, t, used in the function, by $t = x/(2 \times SW)$. If a value of SW = 0.0 is given, the program uses the value for SW which was entered in the par file (or the default). The suggested range of values for s are from 1.3 to 0.7 (possibly 0.5).

zerofill

zerofill < n >

This zerofills the data so that its final size is 2^n times the original size. For data that is not already of size 2^n , the data is first zero padded to the nearest 2^n before further zero filling.

\mathbf{fft}

fft

This does a complex Fourier transform.

phase

phase <phase0> <phase1> <pivot, default=1>

This phases (complex) data with specified parameters <phase0> and <phase1> (both specified in degrees) and <pivot> which assumed by default to be the first point, point 1. The data must be complex.

reduce

reduce

This converts complex data to real data by discarding the imaginary part.

upper

upper <upper bound>

This truncates the data so that the last point is now at <upper bound>. The first point is point 1.

lower

lower <lower bound>

This shifts data so that the origin is now at <lower bound>. The first point is point 1.

range

range <lower bound> <upper bound>

This shifts and truncates data so that the origin is now at <lower bound> and so that the last point is <upper bound>. The first point is point 1.

4.4 Indirect dimension reconstruction

'Reconstruction script' contains the commands for processing the indirect dimensions of NUS acquired data. 'Reconstruction type' provides options for using various different algorithms as well as the FFT for processing the indirect dimensions. 'Extra Params' relevant to the different algorithms are accessed using the dropdown menu. Parameters specific to the different indirect dimensions should be set. The number of available dimensions is governed by the 'ndim' parameter in the 'Data parameters' section.

Reconstruction type

General options

resume

Resumes reconstruction from saved data. The code will attempt to load previous files using default names; if the expected files are not found, the reconstruction will restart afresh. Can be used, for example if max_iter was not set high-enough and further iterations are needed to improve the reconstruction quality.

saveCompleteData

Saves reconstructed time domain data for subsequent use in post-processing software. Use with the post-processing flag 'Input Data useTimeData' (see section 4.5).

record_l2

Write out values of the l2_norm for the reconstruction. The output file is *_l2.out.

record_noise

Write out values of the noise for the reconstruction. The output file is *_noise.out.

Algorithm specific options

The options below are used by some or all of the algorithms. Details are given here. The options available for each algorithm are given in the algorithms section.

max_iter noise l2_stop threshold global_threshold k

stop_flag

\mathbf{FT}

No additional options

IHT

max_iter <default=1000>
noise <noise-level, default=0>
l2_stop <default=0>
threshold <default=0.75>
global_threshold <default=0.01>
stop_flag <None/l2/noise>

\mathbf{IST}

```
max_iter <default=1000>
noise <noise-level, default=0>
l2_stop <default=0>
threshold <default=0.75>
global_threshold <default=0.01>
stop_flag <None/l2/noise>
```

IRLS1

max_iter <1000> noise <noise-level> l2_stop <default=0> k <default=0.015> stop_flag <None/l2/noise>

Dimension commands

scaleT0

scale T0 $<\!\!\rm scaling \ factor\!>$

Scales the first point of the relevant indirect dimension by <scaling factor>. Used to correct the first point of certain experiments. Assumes the data is complex.

conjugate

conjugate

Take the complex conjugate of data. Data is assumed to be complex.

exchange

exchange

Exchange the real and imaginary parts of complex data.

mask

mask_ppmm

Apply ppmm mask (++-) to the relevant indirect dimension i.e. the sign of points 3 and 4 and every third and fourth point equivalently is changed. Needed for some Bruker data.

mask_pm

Apply pm mask (+-) i.e. the sign of every other point is changed. Equivalent to the conjugate command.

extend

extend <no. of complex points>

Extend data by specified number of complex points.

decay

decay <end value>

Multiplies the data by a decaying exponential which is 1 at point 1 and \langle end value \rangle at point n.

sinebell

sinebell <angle>

This multiplies the data by a sine function with the given $\langle angle \rangle$ (specified in degrees) at point 1, and, for n complex points, with angle 180 degrees at point n+1 (*not* point n).

sinebell 2

sinebell2 < angle >

This multiplies the data by a sine function squared with the given $\langle angle \rangle$ (specified in degrees) at point 1, and, if there are n (complex/real) points, with angle 180 degrees at point n+1 (*not* point n).

gaussian_sw

gaussian_sw line broadening (Hz)> <sharpening factor> <spectral width (Hz)>

This multiplies the data by the gaussian $\exp(a + bt + ct^2)$. If LB is the entered line broadening (in Hz) and s is the sharpening factor, then $a = -\ln 2/s^2$, $b = \pi \times LB$ and $c = -(\pi \times LB \times s)^2/4 \ln 2$. Note that LB must be **positive** for normal use, which is the **opposite** of the Bruker convention. The function converts a Lorentzian line (or multiplet) of width FWWH equal to LB to a Gaussian line of width FWWH equal to $s \times LB$. The maximum value of the multiplying function is 1, which is obtained at the fraction of the acquired time $t/T = 2\ln 2/(LB \times \pi \times T \times s^2)$. The spectral width (SW) is needed by the program in order to translate between point number, x, and the time value, t, used in the function, by $t = x/(2 \times SW)$. If a value of SW = 0.0 is given, the program uses the value for SW which was entered in the par file (or the default). The suggested range of values for s are from 1.3 to 0.7 (possibly 0.5).

4.5 Post-processing

Post-processing is available using the Azara processing package. Data input may be timedomain data or frequency domain data from the output of the reconstruction script. This is specified via the flag 'Input Data', using the options *useTimeData* and *useFTData*, respectively. A selection of commonly-used Azara commands can be specified from within the CS-processing script between the flag-lines post-processing script: and end_script. Within this, dimensions can be specified using the Azara nomenclature script_com <dimension>, end_script. An example post-processing script is given below:

```
# Azara post-processing script:
    post_processing_script:
```

```
useFTData
script_com 2
base_poly 6 0
end_script
script_com 3
base_poly 4 0
end_script
script_com 1
base_poly2 8 1 1 400
end_script
end_script
end_script
```

The CS-script uses the commands listed under post_processing_script: to write an Azara script file *_azara.scr, which is then run using the Azara programme (W. Boucher, unpublished; http://www2.ccpn.ac.uk/azara/). This has the advantage that the commands can be re-run or modified at any time by changing the Azara script and calling with process *_azara.scr. The following functions are available from CS-processing using either the GUI or the script-based approach, grouped by category as found in the GUI. A brief description of their usage and syntax is given below, taken from the Azara manual. For more information, see the Azara help section <http://www2.ccpn.ac.uk/azara/azara_docs/process.html>.

Real/complex

complex

complex Specify data is complex.

\mathbf{real}

real Specify data is real.

reduce

reduce

Reduce complex data to real data by throwing away the imaginary part.

Data manipulation

exchange

exchange

Swap the real and imaginary parts of complex data.

conjugate

conjugate

Take the complex conjugate of complex data.

\mathbf{shift}

shift <shift amount>

This shifts the data by the positive amount <shift amount>.

reverse

reverse

This reverses the data.

\mathbf{set}

set <first point> <last point> <value>

Sets each point between <first point> and <last point> (inclusive) to <value>. Applies only to real data. For complex data use **set2**.

$\mathbf{set2}$

set2 <first point> <last point> <real value> <imaginary value>

Sets each point between $\langle \text{first point} \rangle$ and $\langle \text{last point} \rangle$ (inclusive) to $\langle \text{real value} \rangle + i^* \langle \text{imaginary value} \rangle$. Applies only to complex data. For real data use set.

scale

scale <first point> <last point> <value>

Multiply data by <value> from <first point> to <last point> (inclusive). The data must be real. For complex data, use scale2, below.

scale2

scale2 <first point> <last point> <real value> <imaginary value>

Multiply complex data by $\langle real value \rangle + i^* \langle imaginary value \rangle$ between $\langle first point \rangle$ to $\langle last point \rangle$ (inclusive).

$mask_ppmm$

mask_ppmm

Apply ppmm mask (++-) i.e. the sign of points 3 and 4 and every third and fourth point equivalently is changed. Needed for some Bruker data.

mirror_zero

mirror_zero

Mirrors the data for zero dwell time by taking the complex conjugate and reversing the data, excluding the first point. Requires complex data.

mirror_half

mirror_half

Mirrors the data for zero dwell time by taking the complex conjugate and reversing the data, including the first point. Requires complex data.

cycle

cycle <cycle amount>

This cycles the data by <cycle amount>. If the data is complex, <cycle amount> should be specified in terms of complex points.

Data range

lower

lower <lower bound>

This shifts data so that the origin is now at <lower bound>. The first point is point 1.

upper

upper <upper bound>

This truncates data so that its last point is <upper bound>. The first point is point 1.

range

range <lower bound> <upper bound>

This shifts and truncates data so that the origin is now at <lower bound> and so that the last point is <upper bound>. The first point is point 1.

$\underline{\mathbf{fft}}$

zerofill

zerofill < n >

Data is padded with zeros giving a final size of 2^n times the original size where $n \leq 4$

\mathbf{fft}

 fft

Complex Fourier transform. The data must be complex. If the input data is not a power of two, it is zero-padded to the nearest power of two.

\mathbf{ifft}

ifft

Inverse complex Fourier transform. The data must be complex. If the input data is not a power of two, it is zero-padded to the nearest power of two.

\mathbf{hft}

hft

Hilbert Fourier transform. The data must be real. If the input data is not a power of two, it is zero-padded to the nearest power of two.

\mathbf{fftn}

fftn

Normalised complex Fourier transform. The data must be complex. If the input data is not a power of two, it is zero-padded to the nearest power of two.

ifftn

ifftn

Normalised complex inverse Fourier transform. The data must be complex. If the input data is not a power of two, it is zero-padded to the nearest power of two.

Apodisation

decay

decay <end value>

Multiplies the data by a decaying exponential which is 1 at point 1 and $\langle \text{end value} \rangle$ at point n.

decay_sw

decay_sw e broadening (Hz)> <<pre>spectral width (Hz)>

Multiplies the data by a decaying exponential, $\exp(-\pi * LB * t)$, where LB is line broadening (Hz)>. Spectral width is used to convert between points, x, and time values, t, using the function $t = \frac{x}{(2*SW)}$.

sinebell

sinebell < angle >

This multiplies the data by a sine function with the given $\langle angle \rangle$ (specified in degrees) at point 1, and, if there are n (complex/real) points, with angle 180 degrees at point n+1 (*not* point n).

sinebell2

sinebell2 <angle>

This multiplies the data by a sine function squared with the given $\langle angle \rangle$ (specified in degrees) at point 1, and, if there are n (complex/real) points, with angle 180 degrees at point n+1 (*not* point n).

gaussian_sw

gaussian_sw line broadening (Hz)> <sharpening factor> <spectral width (Hz)>

This multiplies the data by the gaussian $\exp(a + bt + ct^2)$. If LB is the entered line broadening (in Hz) and s is the sharpening factor, then $a = -\ln 2/s^2$, $b = \pi \times LB$ and $c = -(\pi \times LB \times s)^2/4 \ln 2$. Note that LB must be **positive** for normal use, which is the **opposite** of the Bruker convention. The function converts a Lorentzian line (or multiplet) of width FWWH equal to LB to a Gaussian line of width FWWH equal to $s \times LB$. The maximum value of the multiplying function is 1, which is obtained at the fraction of the acquired time $t/T = 2\ln 2/(LB \times \pi \times T \times s^2)$. The spectral width (SW) is needed by the program in order to translate between point number, x, and the time value, t, used in the function, by $t = x/(2 \times SW)$. If a value of SW = 0.0 is given, the program uses the value for SW which was entered in the par file (or the default). The suggested range of values for s are from 1.3 to 0.7 (possibly 0.5).

Phase

phase phase <phase0> <phase1>

Phase complex data using the zero order correction, <phase0>, and first order correction, <phase1> (both specified in degrees), with the pivot at point 1, where point 1 is the first point in Azara nomenclature.

phase2

phase2 <phase0> <phase1> <pivot>

Phase complex data using the zero order correction, <phase0>, and first order correction, <phase1> (both specified in degrees), with pivot at point <pivot>, where the first point is point 1.

Baseline correction

base_const

base_const <half width>

This fits the baseline of a spectrum (not fid) by first finding baseline points using a shifting window of half width > half width > and then fitting a constant to those baseline points.

$base_const2$

 $base_const2 < half width > < first point > < last point >$

This fits the baseline of a spectrum (not fid) by first finding baseline points using a shifting window of size <half width> and then fitting a constant to those baseline points. Only the points from <first point> to <last point> (inclusive) are fitted.

base_poly

base_poly <half width> <degree>

This fits the baseline of a spectrum (not fid) by first finding baseline points using a shifting window of size <half width> and then fitting a polynomial of degree (order) <degree> to those baseline points.

$base_poly2$

base_poly2 <half width> <degree> <first point> <last point>

This fits the baseline of a spectrum (not fid) by first finding baseline points using a shifting window of size <half width> and then fitting a polynomial of degree (order) <degree> to those baseline points. Only the points from <first point> to <last point> (inclusive) are fitted.

base_trig

base_trig <half width> <order>

This fits the baseline of a spectrum (not fid) by first finding baseline points using a shifting window of size <half width> and then fitting trig functions of order <order> to those baseline points. In fact, the number of functions used to do the fitting is 2^* <order> + 1, the 2 being because cosines and sines are used, and the 1 being because a constant is also used. This fit only makes sense if the original data set contains the entire recorded spectral width.

base_trig2

base_trig2 <half width> <order> <first point> <last point>

This fits the baseline of a spectrum (not fid) by first finding baseline points using a shifting window of size <half width> and then fitting trig functions of order <order> to those baseline points. In fact, the number of functions used to do the fitting is $2^{*}<$ order> + 1, the 2 being because cosines and sines are used, and the 1 being because a constant is also used. Only the points from <first point> to <last point> (inclusive) are fitted. This fit only makes sense if the original data set contains the entire recorded spectral width.

base_points

base_points

base points file>

This uses the points in the
base points file> to define the baseline points for subsequent baseline correction routines (by default the baseline is determined automatically by the program). There must be a matching end_base_points before the next use of base_points or before a resumption of the automatic baseline determination.

end_base_points

end_base_points

This matches the most recent base_points and must occur before the next use of base_points or before a resumption of automatic baseline determination (the default).

base_subtract

base_subtract

base subtract file>

This subtracts the values given in
base subtract file> from the data. If there are n points then there must be n values in
base subtract file> (in free format).

$base_subtract2$

base_subtract2 <base subtract file> <first point> <last point>

This subtracts the values given in $\langle \text{base subtract file} \rangle$ from the data, from $\langle \text{first point} \rangle$ to $\langle \text{last point} \rangle$ (inclusive). If n = $\langle \text{last point} \rangle - \langle \text{first point} \rangle + 1$, then there must be n values in $\langle \text{base subtract file} \rangle$ (in free format).

Linear prediction

lp_extend

lp_extend <number predicted> <length of sequence> <cutoff>

When <number predicted> is positive, this uses linear prediction to extend the existing data from the end by <number predicted>.

When <number predicted> is negative, this does a linear prediction by <|number predicted|> points inserted at the beginning of the existing data.

Data is assumed to be complex with <number predicted> and <length of sequence> specified in complex points. <cutoff> takes values between 0 and 1; singular values below

cutoff are set to 0. Recommended values for <length of sequence> are $\min(\frac{n}{4} \text{ to } \frac{n}{3}, 15 \text{ to } 20)$ where *n* is the number of complex points prior to linear prediction. The recommended value for <cutoff> is 0.0001. For more details see <http://www2.ccpn.ac.uk/azara/azara_docs/ process.html> and Zhu and Bax, 1992, J. Magn. Reson. 100, 202-207.

lp_forward

lp_forward <number predicted> <number of poles>

Predicts <number predicted> points from the end of the existing data. For complex data, specify in terms of complex points. Real and imaginary points are fitted separately. <number of poles> is an upper bound for the number of oscillators in a given row. A typical value is 5 or 10.

lp_backward

lp_forward <number predicted> <number of poles>

Predicts the first <number predicted> points, overwriting the existing first <number predicted> points. For complex data, specify in terms of complex points. Real and imaginary points are fitted separately. <number of poles> is an upper bound for the number of oscillators in a given row. A typical value is 5 or 10.

lp_first

lp_first <number predicted> <length of sequence> <number of sequences> <cutoff>

Linear prediction of the first <number predicted> points, using <number of sequences> sequences of length <length of sequence>, by overwriting the first <number predicted> points. Specify in terms of complex points. Real and imaginary points are fitted separately. <number of sequence> must be \geq <length of sequence>. Singular value decomposition is used and <cutoff> (between 0 and 1) determines which singular values are significant.

lp_first2

lp_first2 <number predicted> <length of sequence> <number of sequences> <cutoff>

Linear prediction of the first <number predicted> points, using <number of sequences> sequences of length <length of sequence>, by overwriting the first <number predicted> points. Specify in terms of complex points. For complex data, real and imaginary points are fitted as one complex point. <number of sequences> must be \geq <length of sequence>. Singular value decomposition is used and <cutoff> (between 0 and 1) determines which singular values are significant.

lp_last

lp_last <number predicted> <length of sequence> <number of sequences> <cutoff>

Linear prediction of the last <number predicted> points, using <number of sequences> sequences of length <length of sequence>, by extending the existing data. Specify in terms of complex points. Real and imaginary points are fitted separately. <number of sequence>> must be \geq <length of sequence>. Singular value decomposition is used and <cutoff> (between 0 and 1) determines which singular values are significant.

lp_last2

 $lp_last2 < number predicted> < length of sequence> < number of sequence> < cutoff>$

Linear prediction of the last <number predicted> points, using <number of sequences> sequences of length <length of sequence>, by extending the existing data. Specify in terms of complex points. For complex data, real and imaginary points are fitted as one complex point. <number of sequences> must be \geq <length of sequence>. Singular value decomposition is used and <cutoff> (between 0 and 1) determines which singular values are significant.

Chapter 5: Spectrum viewer

A simple spectrum viewer is included with the software to aid visualisation of processed data. The spectrum viewer can be opened in two ways. If the code is run directly from the GUI, the spectrum viewer will open automatically once the reconstruction is completed (assuming post-processing is selected). Alternatively, the viewer can be launched stand-alone using the script graphPlot in the bin directory.

The spectrum viewer reads Azara-style *.spc.par* files (see below) and can display 2, 3 and 4D datasets.

ndim 2 file 102_ft_full_postpro_block.spc dim 1 npts 262 block 64 sw 1000.000 sf 500.000 refppm 0.0000 refpt 1.0000 nuc 1H dim 2 npts 256 block 64 sw 1000.000 sf 500.000 refppm 0.0000 refpt 1.0000 nuc 1H

In each dimension the 'dim', 'npts' and 'block' options are essential. Other parameters contain default values which may be read by Azara plot2. The spectrum viewer is designed for simple visual analysis of the processed spectrum; currently shift referencing is not supported. An example window for a 2D HSQC spectrum is shown in Figure 5.1

1D traces are shown in blue, positive contours in black and negative contours in red. The base level and number of displayed contour levels can be set using the 'Base level' and 'Number of levels' options at the top of the window. In each case a scale factor can be set. For the contour levels, this is multiplicative and determines the change between each contour level for the given number of levels. For the base level, this is also multiplicative and determines how much the base level changes on increasing or decreasing. Some useful commands are given below:



Figure 5.1: 2D spectrum display window.

Command	Function
↑/↓	Increase/decrease the base level by the scale factor
f / d	Scale up/down the 1D traces
shift + left mouse drag	Zoom into the selected region
ctrl + left mouse drag	Zoom into the selected region
Double-click (left-mouse buttom)	Zoom out to the full spectrum size
q	quit viewer
Esc	quit viewer

Table 5.1: Commands for use with the spectrum viewer



Figure 5.2: 3D spectrum display window. 'All' (right-hand side) is selected to stack all the planes in dimension 3 (indicated by 'Stack dims:' on the top toolbar).

For n-dimensional spectrum with n > 2, two dimensions are displayed and the remaining dimensions are 'stacked'. The choice of stacked dimension can be changed using the 'Stack dims' dropdown menu. By default all the planes from the stacked dimension are shown as indicted by the 'all' checkbox on the right hand side. If this is unticked, the slider can then be used to scroll through the stacked dimension(s). The slice selected is shown at the top of the spectrum display window. An example of a 3D display window with all planes in dimension 3 stacked is shown in Figure ??.



Figure 5.3: 3D spectrum display window for a single plane in the first indirect dimension (dimension 2, plane 37). 'All' (right-hand side) is now unticked and the slider used to scroll to the desired slice.

Appendices

Appendix A: Lookup table for Bruker Avance data

Lookup values are given in the main body of the table for a particular (DSPFVS, DECIM) coordinate. Using the lookup value from the table below, groupDelay= $0.5 \times lookup/DECIM$. The first order phase correction applied is $360 \times \langle GRPDLY \rangle$. For more details see the **brukerGroupPhase** command in section 4.3.

DSPFVS	10	11	12	13	20
DECIM					
2	179	184	184	11	-1
3	201	219	219	17	-1
4	533	384	384	23	-1
6	709	602	602	35	-1
8	1097	852	852	47	-1
12	1449	1668	1668	71	-1
16	2225	2312	2292	95	-1
24	2929	3368	3368	143	-1
32	4481	4656	4616	191	-1
48	5889	6768	6768	287	-1
64	8993	9344	9264	383	-1
96	11809	13568	13568	575	-1
128	18017	18560	18560	-1	-1
192	23649	27392	27392	-1	-1
256	36065	36992	36992	-1	-1
384	47329	55040	55040	-1	-1
512	72161	73856	73856	-1	-1
768	94689	110336	110336	-1	-1
1024	144353	147584	147584	-1	-1
1536	189409	220928	220928	-1	-1
2048	288737	295040	295040	-1	-1
2080	-1	-1	-1	-1	282814

Appendix B: Test Data

Test data are contained in the directory *testData*, with processing scripts included to illustrate processing methods.

B.0.1 2D, ¹⁵N HSQC, ubiquitin

This data set was recorded on a Bruker DRX500 spectrometer. Change into the sub-directory 2D. The raw data is contained in the folder 102 in this directory. There are two scripts set-up; 102_ft_full.scr and 102_cs_nus.scr. 102_ft_full.scr is used to demonstrate FT processing of a fully-sampled data set. To load the processing script into the GUI for the fully-sampled HSQC, run the command

cambridgecs -g 102_ft_full.scr

Looking at the GUI display, note that the *Directory* set, $\dots/2D$, is the directory where output files will be stored. Change this on the computer being used by clicking *Select* and choosing the relevant directory. Since the raw data is in the directory 102, the *Raw data directory* parameter is set. This option is available in the *Extra Parameters* drop down menu. An alternative approach would be to set the top level *Directory* to $\dots/2D/102$. In this case the processed data will be stored in the same directory as the raw data. Since the data is fully-sampled, the *Uniform sampling flag* is set, also available from *Extra Paramets*.

The pre-processing script contains commands for FT processing of the direct dimension. Since the data was recorded on a Bruker DRX machine, the parameters DECIM and DSPFVS (found in the Bruker *acqus* file) should be converted to a group delay value and used with the command brukerGroupPhase. Here DECIM = 16 and DSPFVS = 12 and the lookup value (see A) is 2292. This gives brukerGroupPhase = $0.5 \times 2292/16 = 71.625$.

'Reconstruction script' contains commands for processing the indirect dimensions. Here FT is selected from the dropdown menu since the data is fully-sampled and the data can be processed with the fast Fourier transform. In addition, the flag *max_points* must be set (here 256, the total number of points in the indirect dimension), since there is no NUS list for a fully-sampled data set to specify the matrix size.

Post-processing commands are used to add additional baseline corrections using Azara. The input data may be specified here as the FT data i.e. the frequency domain output from the reconstruction.

The button, 'Open script editor' shows the text version of the script, which can be easily adapted to other spectra.

The reconstruction is run by clicking the button *Process current script*. A terminal window displays the progress of the reconstruction with information about the reconstruction parameters and output files. The progress bar on the GUI tracks the progress of the reconstruction (note this does not include the Azara post processing). If Azara post-processing is set, the output spectrum will be displayed in the simple viewer.

The script $102_cs_nus.scr$ demonstrates artificial undersampling of a fully-sampled data set, which may be used in test cases to determine how a particular spectrum responds to different sampling schedules and undersampling factors. The reconstruction is setup largely as above. However, the Uniform sampling flag and max_points are no longer set. Instead two NUS lists are provided; vclist_full and vclist. vclist_full is the Original NUS list which specifies the positions of points in the recorded data. vclist is the NUS list used for undersampling, in this case ~30% sampling following an decaying exponential schedule. The reconstruction is carried out with the IHT algorithm, using the default parameters.

B.0.1.1 NMRPipe

In the folder 2D there is a subfolder, *NMRPipe*, containing NMRPipe processing scripts. The data is prepared using the script *fid.com* which was prepared using the assistance of the 'Bruker' macro:

#!/bin/csh

```
bruk2pipe -in ../102/ser \
  -bad 0.0 -noaswap -DMX -decim 16 -dspfvs 12 -grpdly 0 \
  -xN
                     1024
                           -yN
                                               256
                                                    \backslash
  -xT
                     512
                           -yT
                                               128 \
                          -yMODE
                                          States-TPPI
  -xMODE
                     DQD
                                                        \
               10000.000
                          -ySW
                                         2000.000 \
  -xSW
  -xOBS
                 500.01 -yOBS
                                         50.665588 \
  -xCAR
                   4.679
                           -yCAR
                                               120
                                                    \
  -xLAB
                      HN
                           -yLAB
                                               15N
                                                    ١
                        2
  -ndim
                           -aq2D
                                           States
                                                    \
  -out ./test.fid -verb -ov
```

Direct dimension processing is carried out using the script *nusproc.com*. Importantly, CambridgeCS follows the Azara convention of processing in place along each axis, so there is no need for a transpose statement at the end of the 1D script.

#!/bin/csh

The output from the NMRPipe 1D processing is read into CambridgeCS for reconstruction of the indirect dimensions. In this case, there is no need to setup 'Pre-processing parameters' (pre-processing has been carried out here using NMRPipe). Reconstruction of the indirect dimension is carried out using the commands in the 'Reconstruction parameters' section. Following this, post-processing may be carried out either by selecting Azara post processing commands or using NMRPipe. Here we use NMRPipe:

# G	eneral properties:	
	directory	/home/path/to/CambridgeCS/testData/2D/NMRPipe
	outputFile	102_pipe_cs_2
	dataInput	/home/path/to/CambridgeCS/
		<pre>testData/2D/NMRPipe/test.nus</pre>
	rawDirectory	/102
	nmrPipe	
	uniform_sampling	
	head	512
	direct-dim:	
	npts	261
	end_script	
	indirect-dims:	
	npts	256
	ndim	1
	end_script	
# P	re-processing parameters:	
# R	econstruction parameters:	
	reconstruction_script:	
	reconstruction_flag	FT
	dim 2:	
	zerofill	1
	phase	180.0 -360.0
	mask_ppmm	
	max_points	256
	sinebell	90.0
	end_script	
	end_script	

Azara post processing instructions:

In this example 'uniform sampling' is selected as the data is fully-sampled. In addition the 'nmrPipe' flag is set along with the header flag 'head 512' since all NMRPipe files have a header of length 512 bytes. After reconstruction the processed data will be displayed in the graph viewer window, although note that this does not include any post-processing commands. In addition, a small output script is created and run which converts the data back to NMRPipe format $102_pipe_cs_2.com$:

```
-xFT Freq \
-xMODE Real \
-xSW 2548.828125 \
-xORIG 2661.812500 \
-xOBS 500.010010 \
-xLAB HN \
-yN 512 ∖
-yT 512 ∖
-yFT Freq ∖
-yMODE Real \
-ySW 2000.000000 \
-yORIG 5095.495605 \
-yOBS 50.665588 \
-yLAB 15N \
-ndim 2 \
-out /home/path/to/CambridgeCS/testData/2D/NMRPipe/102_pipe_cs_2.pipe.ft2 -ov
```

The final NMRPipe output file is '102_pipe_cs_2.pipe.ft2'. Further post-processing commands may be added to this or other NMRPipe scripts as needed. Note that if Azara post-processing is selection, the conversion to a NMRPipe output file is also carried out. However, the spectrum displayed in the graph viewer will be the result of the Azara postprocessing.

B.0.2 3D ¹³C NOESY, sensory rhodopsin II (pSRII)

This data set was recorded on a Bruker AV800 spectrometer equipped with a cryoprobe. As for the 2D data set discussed above *Directory* refers to the output directory where the processed data will be saved, whilst *Raw data directory* refers to the directory containing the raw data. The data here is undersampled, using a 25% exponentially decaying NUS schedule. The parameter *nuslist* points to the file *vclist*. The file will be looked for in the *Raw data directory* and this specifies the list of acquired points in the form, where x and y correspond to coordinates for the two indirect dimensions:

x1 y1 x2 y2 x3 y3 ... xn yn

The nuslist option can be found under *Extra params*. This also contains the option *Threads*. This dictates how many CPU threads to use. The default is set to 1. If this value is set to θ , the maximum available will be used. Reconstructions of multidimensional data sets can be speeded up considerably by increasing the number of threads.

Since the data is recorded on and Bruker AV instrument, the GRPDLY parameter in the *acqus* file is used to correct for the first order phase error. This is done automatically with the *brukerGroupPhase* command set to -1.

As the data is undersampled, reconstruction is carried out using one of the compressed sensing algorithms; here iterative hard thresholding (IHT) is used. The stop flag is set to l2; the algorithm compares the l2 norm between successive iterations and will stop when there is no significant change in this value. The IHT may be run with default settings. In this case, the threshold has been increased to 85% of the maximum peak height (0.85). Processing parameters for the two indirect dimensions are set under *dim 2* and *dim 3* and a post-processing script with baseline corrections is also included.

B.0.2.1 NMRPipe

As previously, NMRPipe processing scripts are contained in the subdirectory 'NMRPipe'. For a 3D experiment, each indirect dimension coordinate has four FIDs representing the different quadrature detection components. Thus the y-dimensions has size 4 and the z axis has size determined by the number of coordinates in the *vclist*. The direct dimension processed data is written out as individual 2D planes, which could be used for phasing using for example NMRDraw.

#!/bin/csh

```
bruk2pipe -in ../1734/ser \
  -bad 0.0 -aswap -DMX -decim 2000 -dspfvs 20 -grpdly 67.9862518310547
  -xN
                    1024
                          -yN
                                                4
                                                   -zN
                                                                       480
                                                                            ١
  -xT
                     512
                                                2
                                                   -zT
                                                                       240
                                                                            \
                           -yT
  -xMODE
                     DQD
                           -yMODE
                                         Complex
                                                   -zMODE
                                                                  Complex
                                                                            /
  -xSW
               10000.000
                          -ySW
                                        2000.000
                                                   -zSW
                                                                 3263.708
                                                                            ١
  -xOBS
                 800.134
                          -yOBS
                                         800.134
                                                   -zOBS
                                                                  201.197
                                                                            ١
                          -yCAR
  -xCAR
                   4.773
                                            4.773
                                                   -zCAR
                                                                   22.038
                                                                            ١
  -xLAB
                     1Hx
                           -yLAB
                                               1H
                                                  -zLAB
                                                                       13C
                                                                            ١
  -ndim
                       3
                          -aq2D
                                           States
                                                                            ١
  | pipe2xyz -x -out ./fid.nus/test%03d.fid -verb -ov -to 0
```

Further processing is carried out using the script *csnusproc.com*. Once again, no transpose command is required at the end of the 1D processing script:

```
#!/bin/csh
```

```
xyz2pipe -in fid.nus/test%03d.fid -x \
| nmrPipe -fn SOL \
| nmrPipe
          -fn SP -off 0.5 -end 0.98 -pow 2
| nmrPipe
          -fn ZF -auto
                                                  ١
| nmrPipe
          -fn FT -verb
                                                     \
| nmrPipe
          -fn PS -p0 282 -p1 0 -di
          -fn EXT -left -sw
| nmrPipe
                                        ١
 -ov -out testcs.nus -x
```

Indirect dimension processing is carried out using CambridgeCS with the script 1734_pipe_cs.scr. Again the nmrPipe and head flags are set:

# General properties:	
directory	/home/path/to/CambridgeCS/testData/
	3D/13CNOESY/NMRPipe
outputFile	1734_pipe_cs
dataInput	/home/path/to/CambridgeCS/testData/
	3D/13CNOESY/NMRPipe/testcs.nus
rawDirectory	/1734
nusList	vclist
nmrPipe	540
head	512
direct-dim:	
npts	512
end_script	
indirect-dims:	
npts	1920
ndim	2
end_script	
<pre># Pre-processing parameters:</pre>	
# Reconstruction parameters:	
reconstruction_script:	
reconstruction_flag	IHT
stop_flag	12
dim 2:	
zerofill	2
phase	45.0 0
 mask_ppmm	
sinebell	60.0
end_script	
dim 3.	
zerofill	1
phase	0 0
sinebell	60.0
end_script	
end_script	

Azara post processing instructions:

The NMRPipe script 1734_pipe_cs.com converts the data back to NMRPipe format.

B.0.3 3D SOFAST (Best) [¹⁵N,¹H] TROSY HNCO

The data was recorded on a Bruker DRX500 spectrometer. The reconstruction is set up similar to the 13 C NOESY described above. In this case, 22.7% sampling is used with an

exponentially decaying distribution in the decaying ¹³C dimension (dim 3). The ¹⁵N dimension is recorded with P/N selection; P/N data is handled in the processing by selecting the interlace option from the 'indirect-dims', 'Extra params' drop down menu. *Interlace* is set to 2 to convert dim 2 P/N points to cos/sin points. As for the 2D HSQC discussed above, the first order phase correction is determine from the DECIM and DSPFVS parameters in the *acqus* file. Here DECIM = 16 and DSPFVS = 12 and the lookup value (see A) is 2292. This gives brukerGroupPhase = $0.5 \times 2292/16 = 71.625$.

The reconstruction is carried out using the IHT algorithm. *Global_threshold* is selected from the *Extra Params* dropdown menu and set to 1×10^{-4} . This lowers the minimum threshold for the reconstruction, which may increase the accuracy of the reconstruction, but also increases the reconstruction time and may lead to reconstruction of noise components.

B.0.3.1 NMRPipe

NMRPipe processing proceeds as described above for the 3D NOESY. However, dimension 2 (the first indirect dimension (^{15}N)) is acquired as echo-antiecho data (P-/N-type). In order to avoid the phase-twist line-shape, this data must be converted to cosine-/sine-type data. This is achieved using the Rance-Kay macro provided with NMRPipe which is called using the following command:

```
| nmrPipe -fn MAC -macro $NMRTXT/ranceY.M -noRd -noWr \
```

rance Y.m is used since dimension 2 (y-dimension) needs converting. The full script (*fid.com*) is:

```
#!/bin/csh
```

```
bruk2pipe -in ../550/ser \
  -bad 0.0 -noaswap -DMX -decim 16 -dspfvs 12 -grpdly -1
  -xN
                                              4
                                                                          350
                     1024
                           -yN
                                                    -zN
                                                                               \
                                              2
  -xT
                     512
                           -yT
                                                    -zT
                                                                          175
                                                                               ١
  -xMODE
                     DQD
                           -vMODE
                                             Real
                                                    -zMODE
                                                                       Real
                                                                             \
               10000.000
                                                                  1666.667
  -xSW
                           -ySW
                                         2530.364
                                                    -zSW
                                                                             ١
  -xOBS
                 500.132
                           -yOBS
                                           50.678
                                                    -zOBS
                                                                   125.757
                                                                             ١
  -xCAR
                     4.7
                           -yCAR
                                          120.000
                                                    -zCAR
                                                                   176.000
                                                                             ١
  -xLAB
                      1Hx
                           -yLAB
                                               15N
                                                    -zLAB
                                                                        13C
                                                                             ١
  -ndim
                        3
                           -aq2D
                                           States
                                                                             ١
| nmrPipe -fn MAC -macro $NMRTXT/ranceY.M -noRd -noWr
                                                             \
| pipe2xyz -x -out ./fid/test%03d.fid -verb -ov -to 0
```

Other processing scripts are as previously described, however note that since the Echo-Antiecho data is handled by NMRPipe the *interlace* command should not be set in CambridgeCS.

B.0.4 4D, HCCH NOESY, pSRII

The data was recorded on a Bruker AV800 spectrometer, equipped with cryoprobe. The reconstruction is set up as in previous examples, with the GRPDLY value set to brukerGroupPhase = $0.5 \times 2292/16 = 71.625$.

Reconstruction of a 4D using 12 threads takes around 5 h. For a quick view of the spectrum, it is recommended to change the reconstruction type to \mathbf{FT} . This will allow a

quick overview to check phasing and general spectral appearance before a longer run. It is advised to run longer reconstructions in automatic mode. No user prompts are required and the process can be disconnected from the terminal and run in the backgroup. An example command is given below:

```
nohup nice -n 10 cambridgecs -a 1664_cs.scr &
```

nohup disconnects the process from the terminal ('no hangup' command will be sent if the terminal is closed). *nice -n 10* sets the 'niceness' for the process. Desktop processes receive priority with a value of 0. *cambridgecs -a* runs the process in automatic mode (-*a*) from the command line.

Appendix C: Citation and License

Please cite the following paper when presenting results processed using this software:

Mark J. Bostock, Daniel J. Holland and Daniel Nietlispach (2012), Compressed sensing reconstruction of undersampled 3D NOESY spectra: application to large membrane proteins. J. Biomol NMR, 54:15-32.

CambridgeCS (c) by Mark J. Bostock, Robert Tovey and Daniel Nietlispach, Department of Biochemistry, University of Cambridge, UK.

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Mark J. Bostock, Daniel J. Holland and Daniel Nietlispach (2012), Compressed sensing

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